

Radio-assay of Titanium samples for the LUX Experiment

D. S. Akerib^b, X. Bai^g, S. Bedikian^o, E. Bernard^o, A. Bernstein^e, A. Bradley^b, S. B. Cahn^o, M. C. Carmona-Benitez^b, D. Carr^e, J. J. Chapman^a, Y-D. Chan^d, K. Clark^b, T. Classen^j, T. Coffey^b, S. Dazeley^e, L. de Viveiros^a, M. Dragowsky^b, E. Druszkiewicz^m, C. H. Faham^a, S. Fiorucci^a, R. J. Gaitskell^a, K. R. Gibson^b, C. Hall^l, M. Hanhardt^g, B. Holbrook^j, M. Ihmⁱ, R. G. Jacobsenⁱ, L. Kastens^o, K. Kazkaz^e, R. Lander^j, N. Larsen^o, C. Lee^b, D. Leonard^l, K. Lesko^d, A. Lyashenko^o, D. C. Mallin^g, R. Mannino^h, D. McKinsey^o, D. Meiⁿ, J. Mock^j, M. Morii^c, H. Nelson^k, J. A. Nikkel^o, M. Pangilinan^a, P. D. Parker^o, P. Phelps^b, T. Shutt^b, W. Skulski^m, P. Sorensen^e, J. Spaansⁿ, T. Stiegler^h, R. Svoboda^j, A. Smith^d, M. Sweany^j, M. Szydagis^j, J. Thomson^j, M. Tripathi^{j*}, J. R. Verbus^a, N. Walsh^j, R. Webb^h, J. T. White^h, M. Wlasenko^c, F. L. H. Wolfs^m, M. Woods^j, S. Uvarov^j, C. Zhangⁿ

^aBrown University, Dept. of Physics, 182 Hope St., Providence, RI 02912

^bCase Western Reserve University, Dept. of Physics, 10900 Euclid Ave, Cleveland, OH 44106

^cHarvard University, Dept. of Physics, 17 Oxford St., Cambridge, MA 02138

^dLawrence Berkeley National Laboratory, 1 Cyclotron Rd., Berkeley, CA 94720

^eLawrence Livermore National Laboratory, 7000 East Ave., Livermore, CA 94551

^fMoscow Engineering Physics Institute, 31 Kashirskoe shosse, Moscow 115409

^gSouth Dakota School of Mines and Technology, 501 East St Joseph St., Rapid City, SD 57701

^hTexas A & M University, Dept. of Physics, College Station, TX 77843

ⁱUniversity of California Berkeley, Dept. of Physics, Berkeley, CA 94720-7300

^jUniversity of California Davis, Dept. of Physics, One Shields Ave., Davis, CA 95616

^kUniversity of California Santa Barbara, Dept. of Physics, Santa Barbara, CA 95616

^lUniversity of Maryland, Dept. of Physics, College Park, MD 20742

^mUniversity of Rochester, Dept. of Physics and Astronomy, Rochester, NY 14627

ⁿUniversity of South Dakota, Dept. of Physics, 414E Clark St., Vermillion, SD 57069

^oYale University, Dept. of Physics, 217 Prospect St., New Haven, CT 06511

E-mail: mani@physics.ucdavis.edu

ABSTRACT: We report on the screening of samples of titanium metal for their radio-purity. The screening process described in this work led to the selection of materials used in the construction of the cryostats for the Large Underground Xenon (LUX) dark matter experiment [1]. Our measurements establish titanium as a highly desirable material for low background experiments searching for rare events. The sample with the lowest total long-lived activity was measured to contain <0.25 mBq/kg of ^{238}U , <0.2 mBq/kg of ^{232}Th , and <1.2 mBq/kg of ^{40}K . Measurements of several samples also indicated the presence of short-lived (84 day half life) ^{46}Sc , likely produced cosmogenically via muon initiated (n,p) reactions.

KEYWORDS: Materials for gaseous detectors, Materials for solid-state detectors, Noble-liquid detectors, Gamma detectors.

*Corresponding author.

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1. Introduction

Experiments searching for rare events have to suppress background counts from natural radioactivity, which necessitates the minimization of radio-impurities embedded in the apparatus itself. This is achieved by careful screening of construction materials for the presence of long-lived radio-isotopes, such as ^{238}U , ^{232}Th and ^{40}K . The traditional method for this screening process consists of direct gamma-ray counting at low background facilities, located in laboratories with massive over-burden of rock.

The LUX experiment, which aims to directly detect dark matter, is a two-phase xenon time-projection chamber, containing approximately 350 kg of xenon [1]. In order to sustain such a large mass of xenon in the liquid phase, two concentrically nested cryogenic vessels are employed, with a high vacuum maintained between the inner and outer chambers. The combined mass of these titanium vessels is 230 kg. This article addresses the issue of the choice of material for these cryostats, describes the screening procedure, and presents the results of our measurements of the radio-purity for several titanium samples.

Traditionally, cryostats in large scale physics experiments have been constructed using stainless steel, due to its high tensile strength and relatively low cost; for situations in which radioactivity was critical, copper was preferred [2][3]. Titanium has not been used as the material of choice in any large dark matter experiment to date. The advantages of titanium over copper lie in its superior tensile strength, a somewhat lower cost, and the fact that it can be welded, as opposed to brazing for copper, which can leave micro-holes and allow radioactive contaminants such as radon to enter the experiment. Certain types of stainless steel have been measured to have very low levels of radio-impurities [4] but still quite high compared to copper [5]. Hence, the choice between titanium and copper for the LUX cryostats rested on the radio-assays that are described here.

Titanium metal is available in several grades of chemical purity: the American Society for Testing and Materials (ASTM) lists 39 distinct grades [6], with iron and oxygen content as the most important distinguishing feature. The American Society of Mechanical

Engineers (ASME) has its own nomenclature for categories of titanium metal. Addition of certain impurities, such as oxygen in the form of scrap titanium dioxide, improves the tensile strength of the alloy. Hence, our choice of material consisted of a compromise, because introduction of scrap materials in the processing of a metal carries with it the danger of parasitically introducing long-lived radio-isotopes. We evaluated two grades of titanium in this study, colloquially labeled by vendors in terms of “chemical purity” as CP-1 and CP-2. CP-1 refers to ASTM/ASME B-265 Grade 4 and CP-2 refers to ASTM B-265 Grade 3. However, the CP-1 samples supplied to us were ASTM/ASME B-265 Grade 1.

Tables 1 and 2 show the isotopes of titanium and their cross sections for thermal and fast neutron activation, respectively. The only isotope of concern is the moderately long-lived ^{46}Sc , which is activated in an (n,p) interaction. This process can be initiated by a cosmic ray muon interacting in the environment and producing neutrons via spallation.

Based on estimations of tolerable background rates in LUX, we established a conservative upper limit of 1 mBq/kg for a sample to be considered acceptable. In order to ensure the lowest possible background from radio-isotopes in the cryostat material, several samples of Ti were evaluated in 2008-09. In this article we present the results from this screening process and some follow-up measurements made in 2010-11. The work reported here was carried out at the Low Background Facility at Berkeley Lab, the Oroville Counting Facility in northern California, and the SOLO facility, located at the Soudan mine in northern Minnesota.

2. Method

All the measurements reported in this paper were accomplished through the use of high-resolution gamma ray spectroscopy on samples by direct counting with high-purity germanium (HPGe) detectors located in underground low-background facilities. The detectors used for these measurements, and their Rn-purged shield enclosures, were constructed of low-background materials. Multi-kilogram samples, ranging in weight from 1.5 to 8 kg, were analyzed to achieve the reported results.

We focussed on four benchmark isotopes in our screening of Ti metal samples: ^{238}U , ^{232}Th , ^{40}K , and ^{46}Sc . ^{238}U and ^{232}Th are of particular concern to dark matter experiments for several reasons: a) each isotope decays through a chain in which a large number of high-energy gamma-rays are emitted, b) there is the possibility of generating neutrons via (α ,n) reactions in the surrounding material, and c) there is a small amount of direct generation of neutrons by spontaneous fission in the case of ^{238}U . ^{46}Sc is generated from cosmic ray interactions within the Ti, as described in Sec. 3.1. ^{40}K is a common contaminant in materials, which emits a 1461 keV gamma ray in its decay to ^{40}Ar .

Individual isotope abundances are identified by their respective gamma-ray energies. ^{46}Sc is identified by two $\sim 100\%$ branching ratio gamma-ray lines at 889 and 1120 keV. ^{238}U and ^{232}Th cannot be directly identified by gamma-ray lines from decays of ^{238}U and ^{232}Th nuclei, as gamma-rays from these decays have very low branching ratios and energies. However, decays from isotopes later in the ^{238}U and ^{232}Th decay chains exhibit much higher detectability, and these gamma-rays are typically used to infer the concentration of ^{238}U

and ^{232}Th . The strongest lines for identification of ^{238}U come from ^{214}Pb and ^{214}Bi , and from ^{208}Tl in the case of ^{232}Th . ^{40}K is identified by a single 1461 keV line.

We paid careful attention to the possibility that disequilibria may have been created in the U-series and/or the Th-series as a consequence of chemical processing of the titanium metal. The most important causes for disequilibrium are differences between U/Th and Ra chemistry. Chemical processing during manufacturing can compromise the inference of ^{238}U and ^{232}Th from the late-chain isotopes mentioned above. For instance, a common breakage of equilibrium occurs in the ^{238}U chain for the isotope ^{226}Ra , which is soluble in water and can potentially be introduced or removed in large quantities during certain treatment processes, as can gaseous ^{222}Rn . Because of this, measurements of the ^{235}U decay chain, even though they are less sensitive, are sometimes included as a direct measurement of “early-chain” activity and the ^{238}U content is deduced. The signatures include the 93 keV line from ^{234}Th , the 186 keV line from ^{235}U , and the 1001 keV line from ^{234}Pa . Note that the majority of gamma-rays contributing significantly to electromagnetic backgrounds in the detector are generated from late-chain decays. Hence, measurements of early-chain concentrations tend to result in very weak upper limits, and direct measurement of late-chain activity is typically considered to be of greater importance for background estimation.

2.1 Facilities

The Low Background Facility (LBF) at LBNL was established by A.R. Smith in the late 1950s for quantifying the stray radiation field of the newly operational Bevatron. Since the early 1980s, the LBF has been increasingly utilized to select low-activity construction materials for use in experiments searching for rare events. It features a 4π -shielded room with concrete walls (4-6 ft thick) made from about 500 tons of selected low-radioactivity serpentine rock ($\text{Mg}_6\text{Si}_{10}\text{OH}_8$) concrete. The main detector, a 115% N-type Low Background HPGe Detector, is housed in an outer Pb and an inner Cu (OFHC) shielding layer.

In addition, data were collected at an offsite facility located under the Oroville dam in northern California, which provides 450 meters water equivalent (mwe) of shielding. An 85% P-Type Ultra-low Background HPGe detector is housed within low-activity Pb and Cu shields and continuously flushed with nitrogen. It has now lived underground for about 8 years and has routinely provided sensitivities at levels of <1 mBq/kg $^{238}\text{U}/^{232}\text{Th}$ for late-chain measurements.

Data were also taken with a 0.6 kg HPGe detector at the SOLO facility [7] located at the Soudan Underground Laboratory, and operated by the LUX experiment. The Soudan lab boasts 2200 mwe of overhead shielding. The detector is housed in a lead shield >30 cm thick, with the inner 5 cm comprised of ancient lead with ^{210}Pb activity measured below 50 mBq/kg. The detector offers sensitivities at the level of ~ 1 mBq/kg $^{238}\text{U}/^{232}\text{Th}$ for late-chain measurements.

3. Results of Radio-assays

Nine different samples of Ti (seven samples of CP-1 and two samples of CP-2), obtained from various distributors, were counted at the three different facilities between February

2008 and August 2011. Samples Ti1 - Ti7 are plate samples, while Ti8 is a sheet, and Ti9 is a sample of Ti welding wire. Sample descriptions are listed in Table 3.

Counting results for CP-1 samples are listed in Table 4. All quoted errors are at $\pm 1\sigma$, while upper limits are at 90% CL. The (e) and (l) labels refer to early- and late-chain measurements, as described in Sec 2. Ti1, Ti4, Ti5, and Ti6 are plate samples of similar thicknesses and consistently show only upper limits for radioactive contaminations, except in the case of Ti6, for which an early-chain estimate of 6.2 ± 1.2 mBq/kg was made from measuring the 186 keV line from ^{235}U decays. This sample has the lowest measured ^{238}U upper limit at <0.19 mBq/kg, and was ultimately used for the construction of the body of the LUX cryostats. While samples Ti1, Ti4, and Ti5 did not have a positive count for early-chain ^{238}U , the low late-chain counts of Ti6 make it the superior material.

Ti7 is a thick plate sample which was used to fabricate the sealing flanges and end-caps for the LUX cryostats. It has the lowest measured ^{232}Th upper limit at <0.2 mBq/kg and ^{238}U upper limit at <0.25 mBq/kg. Sample Ti8 was used in small quantities for fabricating peripheral parts. Hence, even though it shows positive measurements for late-chain decays, the total mass used in LUX was negligible compared to the bulk materials (Ti6 and Ti7). Ti9, which is the welding wire used in construction, also has a positive signature for both early- and late-chain decays. The Ti9 late-chain signature is more than twice that measured for Ti8, more than twice the highest upper limit (made for sample Ti4), and $\times 28$ higher than the lowest upper limit placed. Again, negligible quantities of this material were employed.

Results for CP-2 are listed in Table 5. Sample Ti3 was measured with a large positive ^{238}U contamination, at 37 ± 12 mBq/kg. We were not able to place any strong limit on ^{238}U contamination of sample Ti2, as the baseline activity was significantly raised from Compton scattering of 2.6 MeV gamma-ray associated with the highly abundant ^{232}Th in this sample.

Comparison of counting results for CP-1 and CP-2 reveals that CP-1 has a much lower average ^{232}Th content than CP-2. A difference in ^{238}U contamination is also seen between CP-1 and CP-2. Measurements for CP-1 ^{238}U contamination yield only upper limits for four out of seven samples, with a mean positive measurement of 9.9 ± 1.8 mBq/kg (early) for the remaining three. Late-chain activities for CP-1 are significantly lower, with only samples Ti8 and Ti9 indicating positive detection at an average of 2.1 mBq/kg.

No positive signatures of ^{40}K were identified, except in the case of the wire sample Ti9. Given the positive identification of ^{238}U / ^{232}Th at levels higher than typical for other CP-1 materials (particularly in the case of ^{232}Th), as well as the presence of ^{40}K , it is likely that this sample was contaminated during chemical processing. It should be noted that the wires were not deliberately thoriated, as is the practice for welding applications.

3.1 Measurements of Cosmogenic Activation

Cosmogenic activation, while possible at sea level, is enhanced during air shipment. The primary cosmogenic activation product in Ti is ^{46}Sc , with a half-life of ~ 84 days. It decays by emitting two ~ 1 MeV gamma-rays (889 and 1120 keV), along with a beta with an endpoint energy of 356 keV. The cross section for the dominant production via fast

neutron activation, $^{46}\text{Ti}(n,p)^{46}\text{Sc}$, is $\sigma \sim 0.25$ barn for neutron energies in the range 5-20 MeV. Neutrons in this energy range are predominantly generated from muon spallation in nearby materials.

Positive measurements of ^{46}Sc were made for several of the Ti samples, as shown in Tables 4 and 5. In some cases, upper limits were extracted, but not for each sample. The measured activities range from below 0.19 mBq/kg to values as high as 23 mBq/kg (Ti6). We believe that this variation is due to their history of transportation, which was not monitored prior to their arrival in our labs. A subset of Ti6, listed as Ti6-A, was counted again after a gap of about 3 years and no peaks due to ^{46}Sc were visible. Similarly, Ti3 underwent a count after about a two-year gap, and no ^{46}Sc lines were observed. This leads us to conclude that cosmogenic ^{46}Sc will not be an issue for titanium stored at locations near sea level (in this case, Davis, CA). However, the surface lab at Lead, SD is situated about 1,500 m above sea level, and presents possible concerns for cosmogenic activity.

We did a more detailed study of ^{46}Sc activation using sample Ti5. It was initially counted at SOLO after being underground for ~ 300 days, yielding <0.35 mBq/kg of ^{46}Sc . It was then sent to Sanford Lab for a period of ~ 6 months, after which it was shipped by ground back to SOLO for counting. Counting after shipment yielded an activity of 4.4 ± 0.3 mBq/kg, averaged over a counting period of 14 days. Ideally, we should expose a sample to cosmic rays for a period about 5 times larger than the half-life of the activation product, in order to achieve equilibrium levels. However, this measurement provides us with a good estimate for the equilibrium level of ^{46}Sc that can be expected in all titanium parts that were stored at the Sanford Lab for a period of 6 months or more.

4. Conclusions

From the nine Ti samples screened in the LUX materials selection program, the CP-1 plate samples were found to have a significantly reduced ^{232}Th and ^{238}U content compared to the CP-2 samples. The CP-1 sheet and wire samples were not significantly different from CP-2 plate samples. The bulk materials used in the construction of the LUX cryostat were measured to have total activities well below 1 mBq/kg, our established tolerance level. Within our limited data set, it is difficult to draw definitive conclusions about radioactive properties of various types of titanium metal. However, it is evident that ASTM/ASME B-265 Grade 4 (CP-1) titanium is a very promising material for low background experiments. Prolonged exposure of titanium to cosmic rays at sea level is not of concern, but storage at high altitudes causes a build-up of ^{46}Sc levels.

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References

- [1] The LUX Collaboration, Akerib *et al.*, “The Large Underground Xenon (LUX) Experiment”, In preparation.
- [2] <http://arxiv.org/abs/astro-ph/0605500> (The ZEPLIN Experiment)
- [3] <http://arxiv.org/pdf/1005.0380> (The XENON100 Experiment)
- [4] <http://radiopurity.in2p3.fr/search.php?Material=Steel>
- [5] <http://radiopurity.in2p3.fr/search.php?Material=copper>
- [6] <http://www.astm.org/Standards/B381.htm>
- [7] Attisha *et al.*, “Soudan Low Background Counting Facility (SOLO)”, *Proceedings of the Topical Workshop on Low Radioactivity Techniques*, 2005, published by AIP.

Table 1. A table of stable isotopes of titanium and their activation cross sections when exposed to thermal neutrons.

Isotope	Abundance (%)	Cross Section (barns)	Product	Half-life (min)
^{46}Ti	8.0	0.59	^{47}Ti	stable
^{47}Ti	7.3	1.7	^{48}Ti	stable
^{48}Ti	73.8	7.84	^{49}Ti	stable
^{49}Ti	5.5	2.2	^{50}Ti	stable
^{50}Ti	5.4	0.179	^{51}Ti	5.77

Table 2. A table of stable isotopes of titanium and their activation cross sections via (n,p) reactions at a neutron energy of 5 MeV.

Isotope	Abundance (%)	Cross Section at 5 MeV (mb)	Product	Half-life (hrs)
^{46}Ti	8.0	90.47	^{46}Sc	2010.72
^{47}Ti	7.3	61.40	^{47}Sc	80.38
^{48}Ti	73.8	0.14	^{48}Sc	43.67
^{49}Ti	5.5	5.11	^{49}Sc	0.95

Table 3. Descriptions of all Ti samples used in the radio-assay for this work. Sample IDs are matched to counting results given in the tables below.

ID	Grade	Type	Dimensions (inches)	Quantity	Total Mass (kg)
Ti1	CP-1	Plate	2.5×6×0.375	4	1.873
Ti2	CP-2	Plate	4×6×0.188	20	7.55
Ti3	CP-2	Plate	1.25×6×0.358	8	1.546
Ti4	CP-1	Plate	4×6×0.188	18	7.34
Ti5	CP-1	Plate	4×6×0.188	16	6.51
Ti6	CP-1	Plate	4×6×0.188	20	7.978
Ti6-A	CP-1	Plate	4×6×0.188	9	3.586
Ti7	CP-1	Plate	1.875×6×1	8	7.201
Ti8	CP-1	Sheet	4×6×0.0625	40	4.399
Ti9	CP-1	Wire	0.0625 dia. x 6 long	780	4.189

Table 4. Radioactivity assay results for 7 different CP-1 Ti samples. All reported errors are at $\pm 1\sigma$; all upper limits are at 90% CL. The samples Ti6 and Ti7 were used as the bulk materials for the construction of LUX cryostats. Samples Ti8 and Ti9 were used in small quantities for fabricating peripheral parts and for welding, respectively.

ID	Facility	Counting Period	^{238}U (mBq/kg)	^{232}Th (mBq/kg)	^{40}K (mBq/kg)	^{46}Sc (mBq/kg)
Ti1	Oroville	Feb 2008	<2.5	<1.6	<6.2	4.8
Ti4	LBF	Jun 2008	<3.8	<2.8	<9.3	3.0
	LBF	Oct 2010	<6.3 (e) <1.3 (l)	<1.2	<9.3	
Ti5	SOLO	2008	<2.5 (l)	<0.8 (l)		<0.35
Ti6	Oroville	Sep 2008	6.2 $\pm$ 1.2 (e) <0.19 (l)	<0.24	<0.93	23 $\pm$ 1
Ti6-A	LBF	Aug 2011	<11.3 (e) <0.25 (l)	<1.6	<9.3	
Ti7	Oroville	Jan 2009	<0.25	<0.2	<1.2	2.5
Ti8	Oroville	Sep 2009	10 $\pm$ 3.8 (e) 2.4 $\pm$ 0.5 (l)	<0.8 (e) 2.8 $\pm$ 0.4 (l)	<1.9	6 $\pm$ 1
Ti9	Oroville	Sep 2009	15 $\pm$ 3.8 (e) 1.9 $\pm$ 0.4 (l)	8.1 $\pm$ 1.2 (e) 5.7 $\pm$ 0.8 (l)	7.4 $\pm$ 1.9	

Table 5. Radioactivity assay results for two different CP-2 Ti samples. All reported errors are at $\pm 1\sigma$; all upper limits are at 90% CL.

ID	Facility	Counting Period	^{238}U (mBq/kg)	^{232}Th (mBq/kg)	^{40}K (mBq/kg)	^{46}Sc (mBq/kg)
Ti2	SOLO	2008	<140 (l)	70 $\pm$ 5 (l)	<9.3	
Ti3	SOLO	2008	<27	<22		<23
	Oroville	Nov 2010	37 $\pm$ 12 (e) 7.4 $\pm$ 1.2 (l)	5.7 $\pm$ 2 (e) 16 $\pm$ 2 (l)	<5	