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## Introduction to PIC Codes and Astra

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The ASTRA code+documentation from Klaus Floettmann: http://www.desy.de/~mpyflo/

> "Computer Simulation Using Particles" R W Hockney, J W Eastwood

Robert Ryne, "Computational Methods in Accelerator Physics" http://uspas.fnal.gov/materials/09UNM/ComputationalMethods.pdf



- Why simulations?
- Commonly-used simulation tools
- Particle-in-cell algorithm overview
- Numerical artifacts and convergence
- Introduction to ASTRA
  - 1) generating the initial beam
  - 2) the main ASTRA input file
  - 3) diagnostic output



#### "The purpose of computation is insight, not numbers." Richard Hamming

Simulations quantify the expected performance of proposed machines to produce high confidence that systems will work as intended/promised to funding agencies.

Simulations build intuition about machine performance before experiments begin.

- Offer full, non-intrusive diagnostics at nearly any location in the machine
- Effects can be turned on/off to isolate

Simulations allow analysis of more realistic situations than analytically tractable.

- Detailed and complex geometries
- Non-ideal beam distributions
- Nonlinear effects

Simulations are cheaper and faster than running the experiment, and can exploit dramatic ongoing improvements in computational hardware and software.

The highest understanding and confidence is achieved when results from analytical theory, numerical simulation, and experiment all converge.



A number of codes have been developed over the years that are heavily used in the design and optimization of high-brightness electron injectors. These codes must be capable of simulating electron beams in the presence of intense nonlinear space charge forces, which play a critical role at low energy.

An incomplete list of 'popular' codes for injector simulation include (alphabetical order):

- ASTRA: free downloadable at <a href="http://www.desy.de/~mpyflo/">http://www.desy.de/~mpyflo/</a>
- GPT: commercial (<u>http://www.pulsar.nl/gpt/</u>)
- HOMDYN: free, contact Massimo Ferrario (Massimo.Ferrario@Inf.infn.it)
- IMPACT-T: free, contact Ji Qiang (jqiang@lbl.gov)
- PARMELA: free, export limitations apply. <u>http://laacg1.lanl.gov/laacg/services/</u>

All of the above are *particle* codes, with the exception of HOMDYN, which is an *envelope* code. Particle codes represent the dynamics of the physical beam by tracking the positions and momenta of a large number  $N_p$  of simulation particles or *macroparticles*.





#### **Field Solution**



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Three numerical parameters are critical for any PIC simulation:
1)time step: Δt
2)grid cell size: Η
3)number of simulation particles: N<sub>p</sub>

Even if the external fields throughout the machine are precisely known, the use of discretized time  $\Delta t$  and space H will limit the resolution of physics that can be captured accurately by the simulation, as well as introducing unphysical effects.

Typically  $N_p \ll N_e$ , where  $N_e$  is the physical number of electrons in the beam. This artificially increases the fluctuations of self-fields experienced by the particles. In injector simulations, these effects primarily result in additional growth of the beam emittances and energy spread not present in the physical electron beam.

It is therefore important to check the convergence of the simulation output with respect to the three parameters above in the limit as:

$$\Delta t \longrightarrow 0$$
,  $H \longrightarrow 0$ ,  $N_p \longrightarrow N_e$ .



#### A Space charge TRacking Algorithm

- **1)** Developed at DESY for photoinjector simulations
- **2)** Serial and parallel implementations (will focus on serial)
- 3) A collection of different programs/utilities
  - 1) generator Generates the initial distribution from the distribution input file
  - 2) astra Propagates the particles according to the astra input file
  - 3) fieldplot Plots the external and space charge fields
  - 4) postpro Calculates and plots properties of the final distribution
  - 5) lineplot Plots the evolution of beam properties along the beamline

Both 2-D (cylindrical coordinates R-Z for axisymmetric systems) and 3-D field solvers are available. We will be using the 2-D solver only.



#### Calculation of external fields

- For RF fields input the longitudinal on-axis electric field  $E_z(z,r=0)$
- For solenoids input the longitudinal on-axis magnetic field B<sub>z</sub>(z,r=0)
- Fields off-axis are determined from Maxwell's equations using Taylor series about the axis.

#### **Calculation of space charge fields (2-D)**

- Use a cylindrical grid with Nlong\_n longitudinal and Nrad transverse cells.
- Transform to the beam rest frame, assume constant charge density within each cell and integrate to get the SC force.

#### Particle pusher

- Using the internal and external fields, the individual particles are propagated using a Runge-Kutta 4<sup>th</sup> order integrator.
- The minimum and maximum time steps for the integrator are user defined, and should be checked for convergence.



#### Use "generator" on a distribution input file with suffix "--.in"



Format of gonarated	Number	1	2	3	4	5	6	7	8	9	10
Pormal of generaled	Parameter	Х	у	Z	рх	ру	pz	clock	macro	particle	status
particle burich									charge	index	flag
	Unit	m	m	m	eV/c	eV/c	eV/c	ns	nC		

# Astra assumes a planar cathode and includes image charge forces (when the cylindrical 2-D algorithm is used).

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### **Structure of the Main Input File**

This is a file with suffix "in"	&NEWRUN Head=' Example of ASTRA users manual' PUN=1						
Initial Beam	NON=1 Distribution = 'Example.ini', Xoff=0.0, Yoff=0.0, TRACK_ALL=T, Auto_phase=T						
	H_max=0.001, H_min=0.00						
Diagnostics	&OUTPUT ZSTART=0.0, ZSTOP=1.5 Zemit=500, Zphase=1 RefS=T EmitS=T, PhaseS=T /						
Field Solver							
Space charge is off. Set LSPCH=T to include —— space charge forces!	<pre>&amp;CHARGE    LSPCH=F    Nrad=10, Cell_var=2.0, Nlong_in=10   Number of radial and         longitudinal grid points         Max_Scale=0.05    /</pre>						
Layout of elements	&CAVITY LEField=T, File_Efield(1)='3_cell_L-Band.dat', C_pos(1)=0.3 Nue(1)=1.3, MaxE(1)=40.0, Phi(1)=0.0, /						
This value is added to the z-coordinate of points in the on-axis field input file. If the input file data is symmetric around zero, then this value gives the distance of the element's midpoint from the cathode.	<pre>&amp;SOLENOID   LBField=T,   File_Bfield(1)='Solenoid.dat', S_pos(1)=1,2   MaxB(1)=0.35, S_smooth(1)=10 /</pre>						



#### **Output particle distribution: Name.zpos.001**

Column format is the same as that used for the initial particle distribution.

Number	1	2	3	4	5	6	7	8	9	10
Parameter	Х	У	Z	px	ру	pz	clock	macro charge	particle index	status flag
Unit	m	m	m	eV/c	eV/c	eV/c	ns	nC		

#### Output rms moments: Name.Xemit.001, Name.Yemit.001, Name.Zemit.001

Contain the evolution of the bunch centroid, rms size, and emittance along the beamline.

Note: the emittance calculation uses canonical rather than mechanical momentum.

$$\langle x \rangle = \frac{1}{N} \sum_{j=1}^{N} x_j \qquad x' = p_x/p_z$$

$$\epsilon_{nx}^2 = \frac{1}{m^2 c^2} \left( \langle x^2 \rangle \langle p_x^2 \rangle - \langle x p_x \rangle^2 \right)$$

#### Columns

Xemit	z	t	x <sub>avr</sub>	x <sub>rms</sub>	x' <sub>rms</sub>	$\epsilon_{x,norm}$	x·x' <sub>avr</sub>
	m	ns	mm	mm	mrad	$\pi$ mrad mm	mrad
Yemit	z	t	y <sub>avr</sub>	y <sub>rms</sub>	y' <sub>rms</sub>	ε <sub>y,norm</sub>	y·y' <sub>avr</sub>
	m	ns	mm	mm	mrad	π mrad mm	mrad
Zemit	z	t	E <sub>kin</sub>	Z <sub>rms</sub>	ΔE <sub>rms</sub>	ε <sub>z,norm</sub>	z·E' <sub>avr</sub>
	m	ns	Mev	mm	kev	π keV mm	keV

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